

# Applications of multi-beam CLEAN

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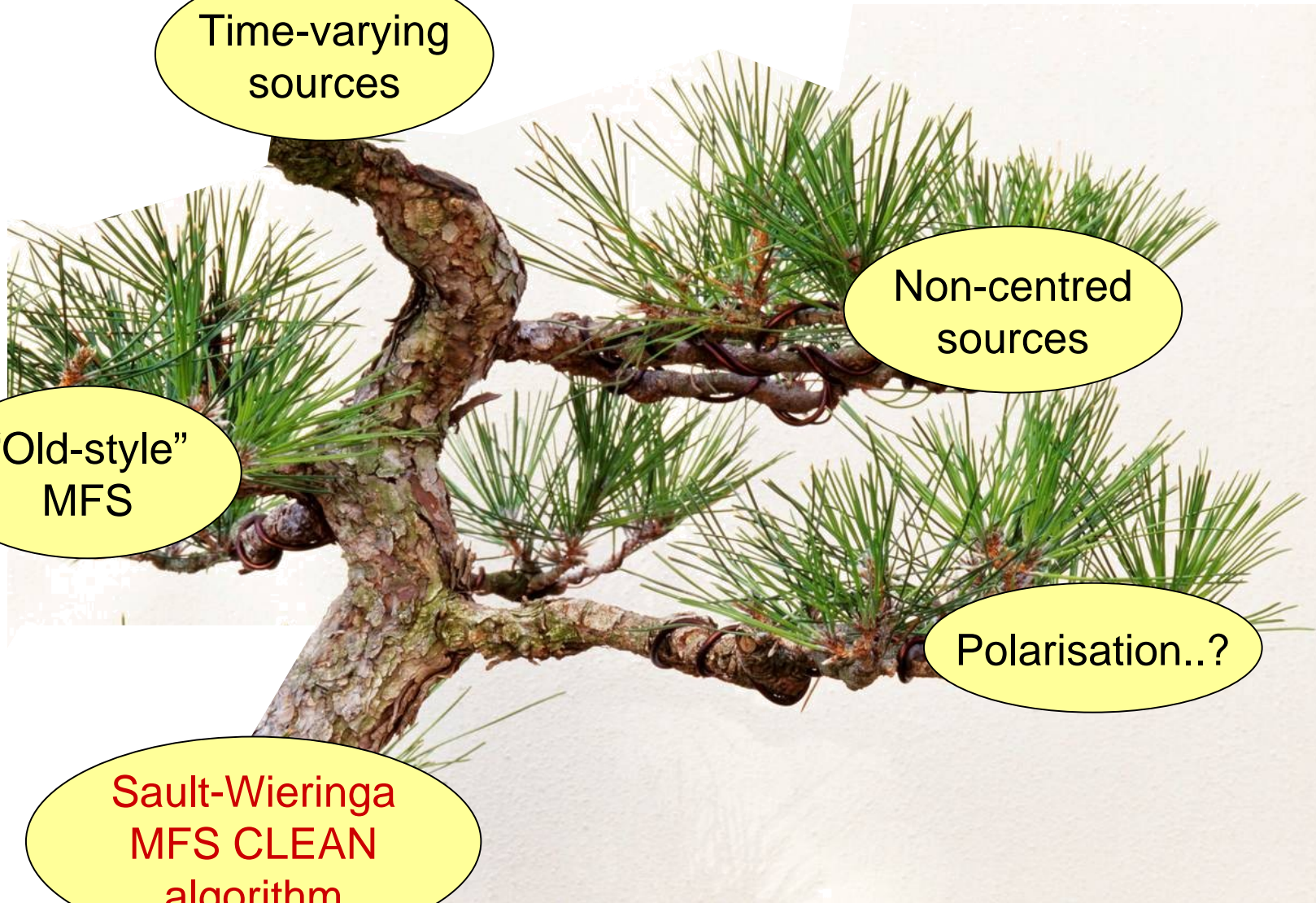
Time-varying  
sources

Non-centred  
sources

“Old-style”  
MFS

Polarisation..?

Sault-Wieringa  
MFS CLEAN  
algorithm



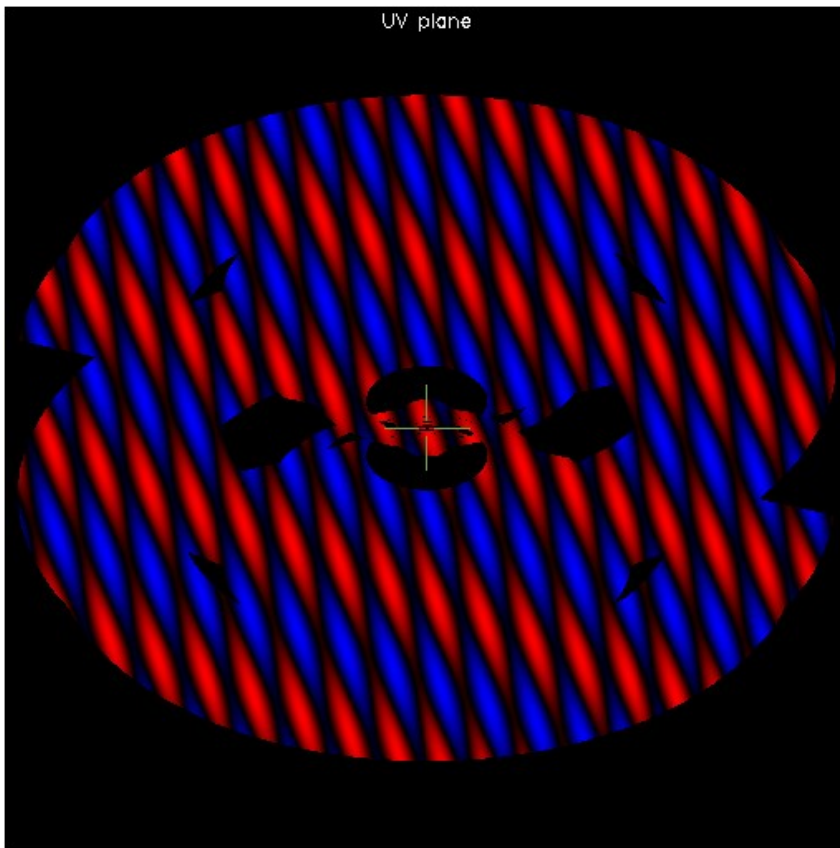
# Review of the S-W algorithm

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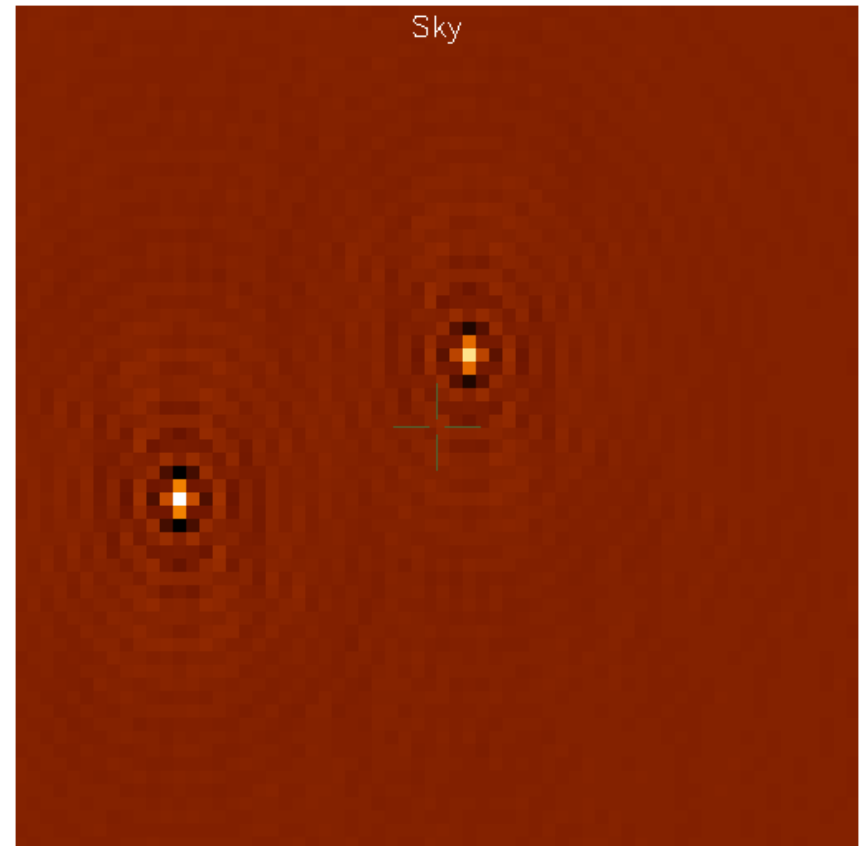
# Review of the S-W algorithm

Wide-band frequency synthesis where all sources have flat spectra:  
standard CLEAN works perfectly\*.

\*well, *almost* perfectly.



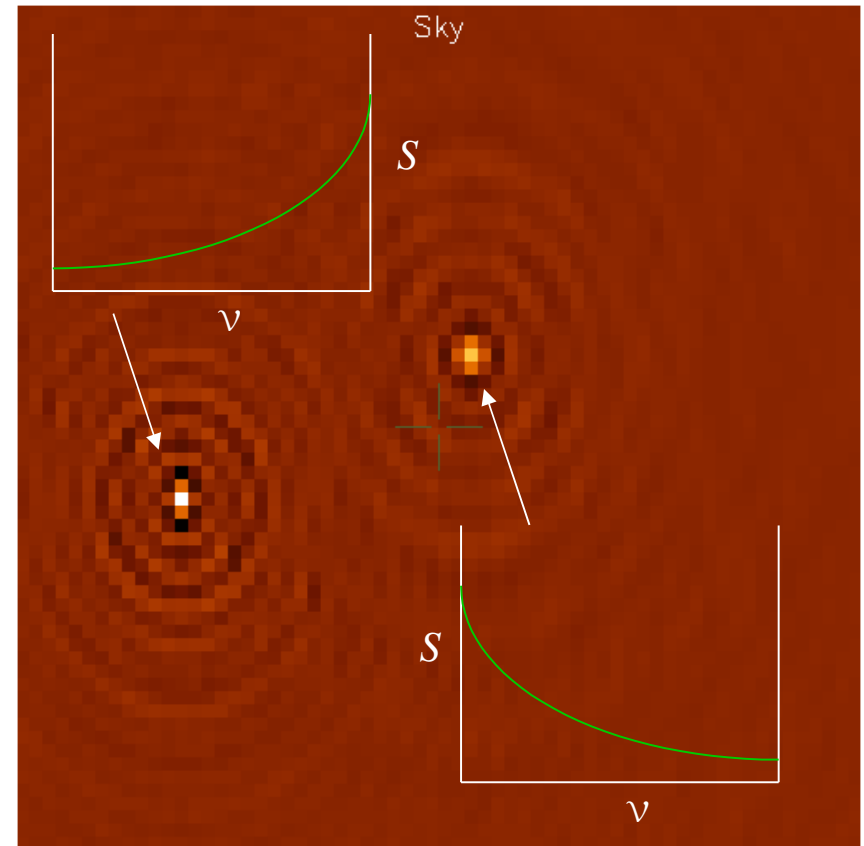
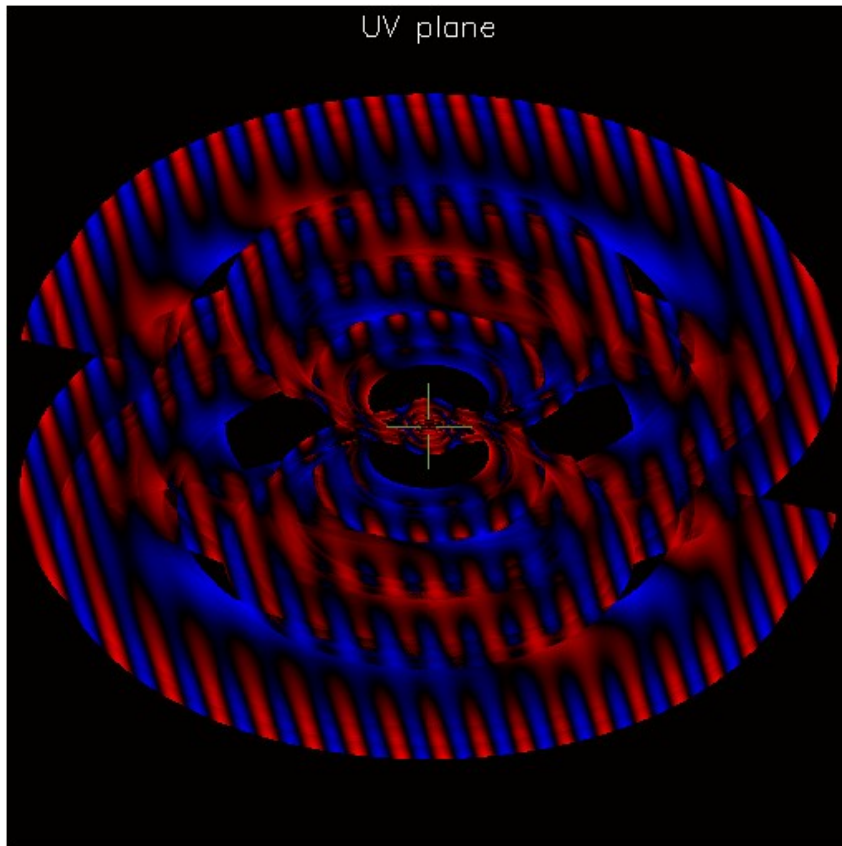
Simulated e-Merlin visibilities, 5 – 7 GHz,  
 $\delta = +35^\circ$ , uniform weighting.



Dirty image.

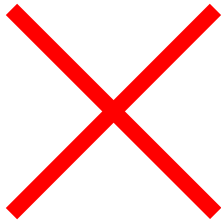
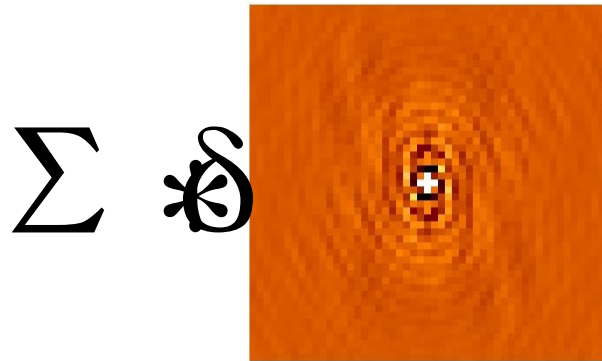
# Review of the S-W algorithm

But, in the real world, different sources usually have different spectra.  
In this case the CLEAN assumption breaks down.

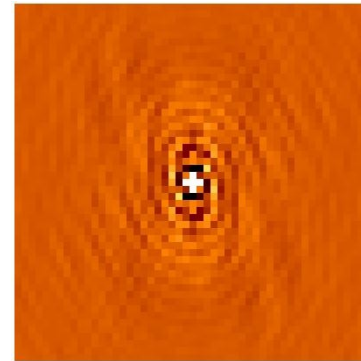


# Review of the S-W algorithm

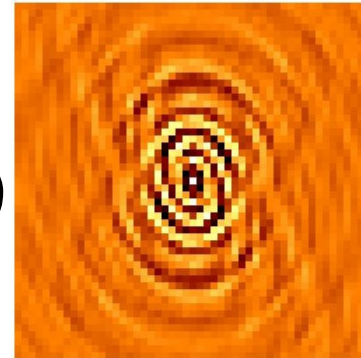
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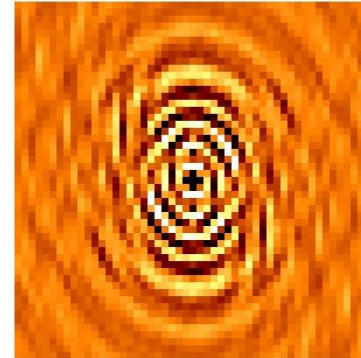
$\Sigma * \delta$



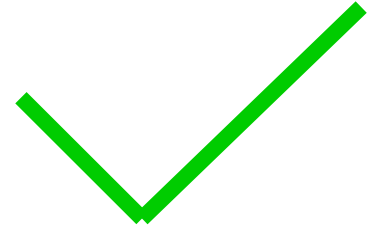
+



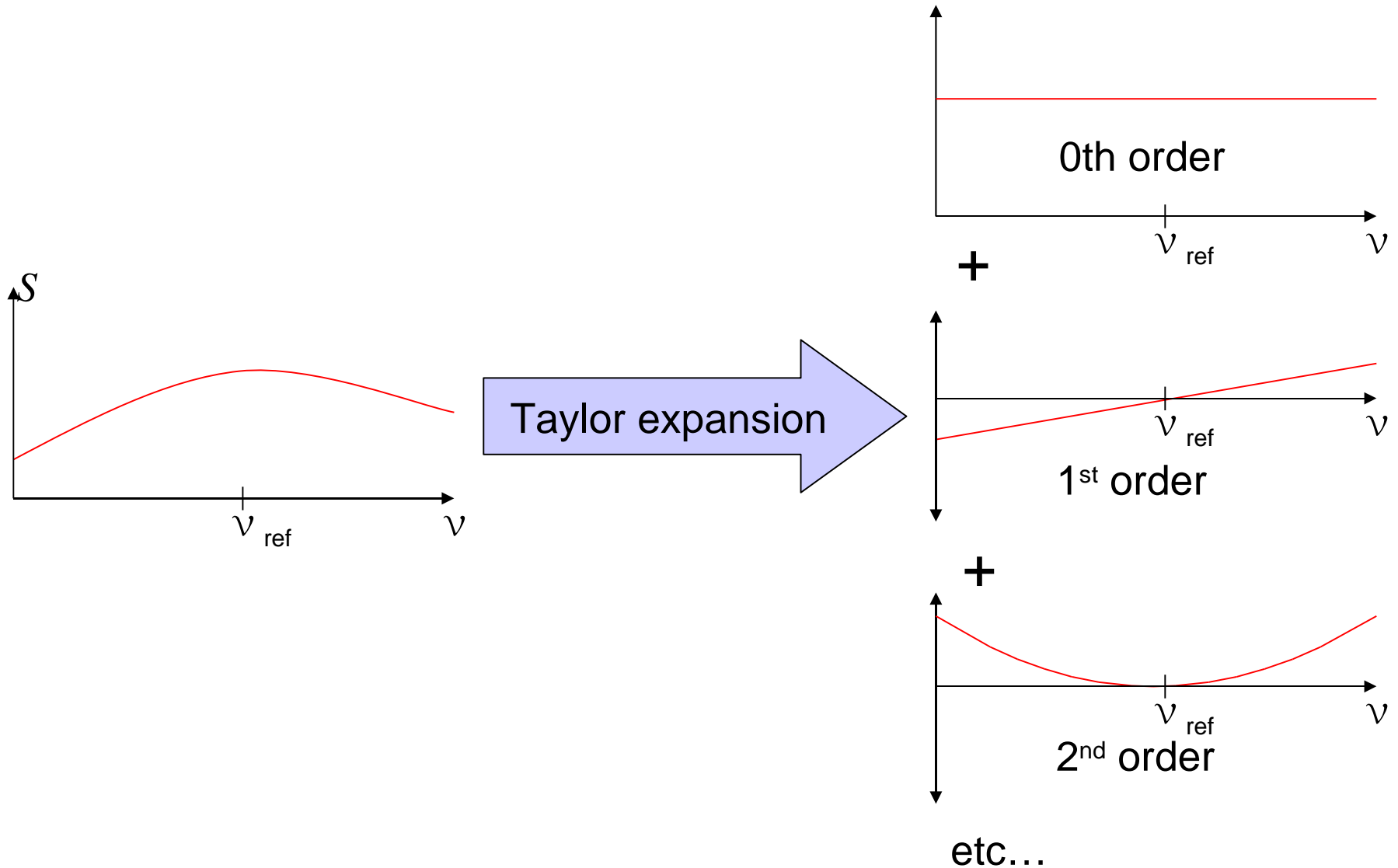
+



+ ...



# Review of the S-W algorithm

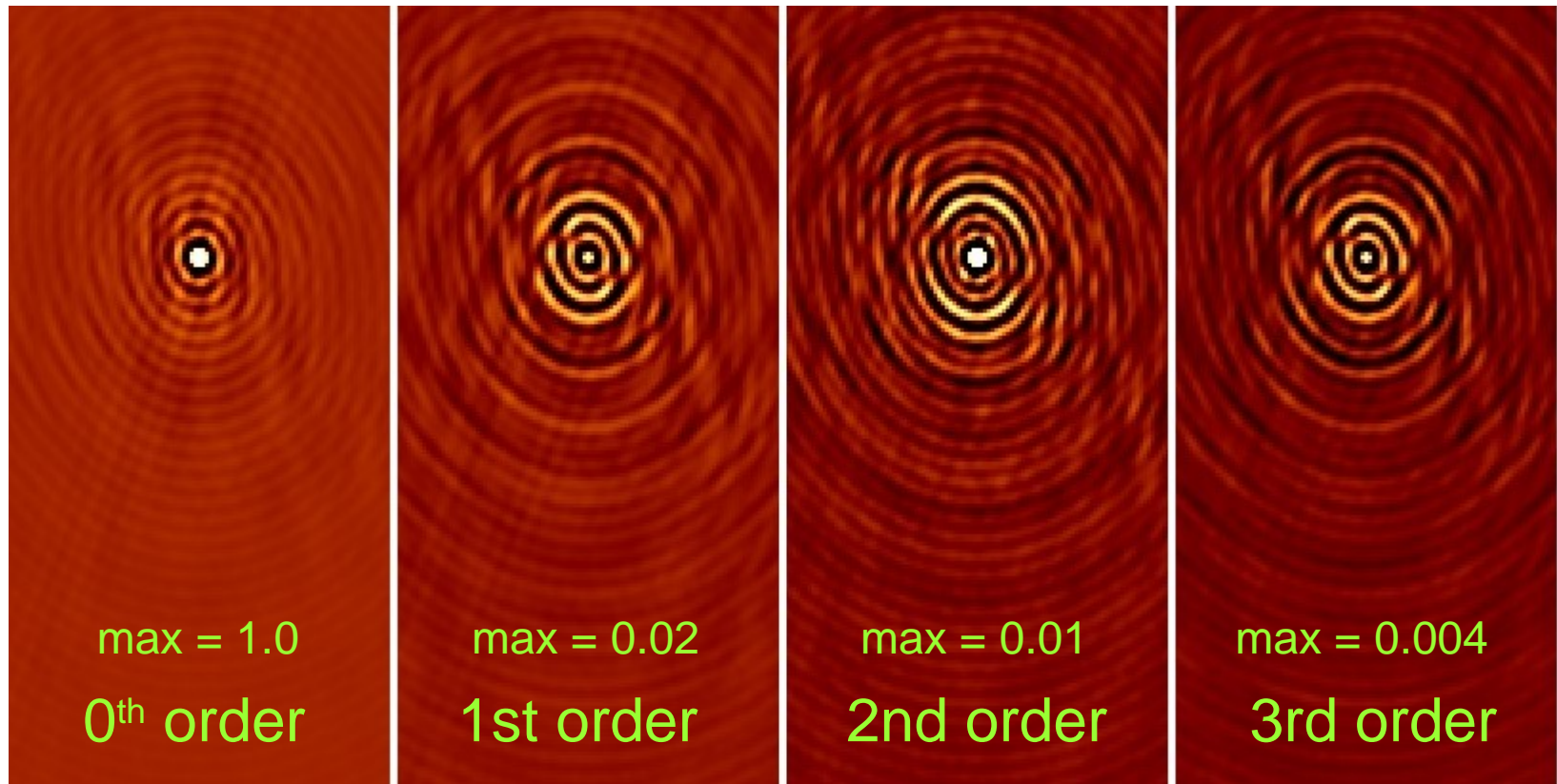




# Review of the S-W algorithm

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Taylor-series beams for a typical e-Merlin observation.



# Fundamental aim of this talk:

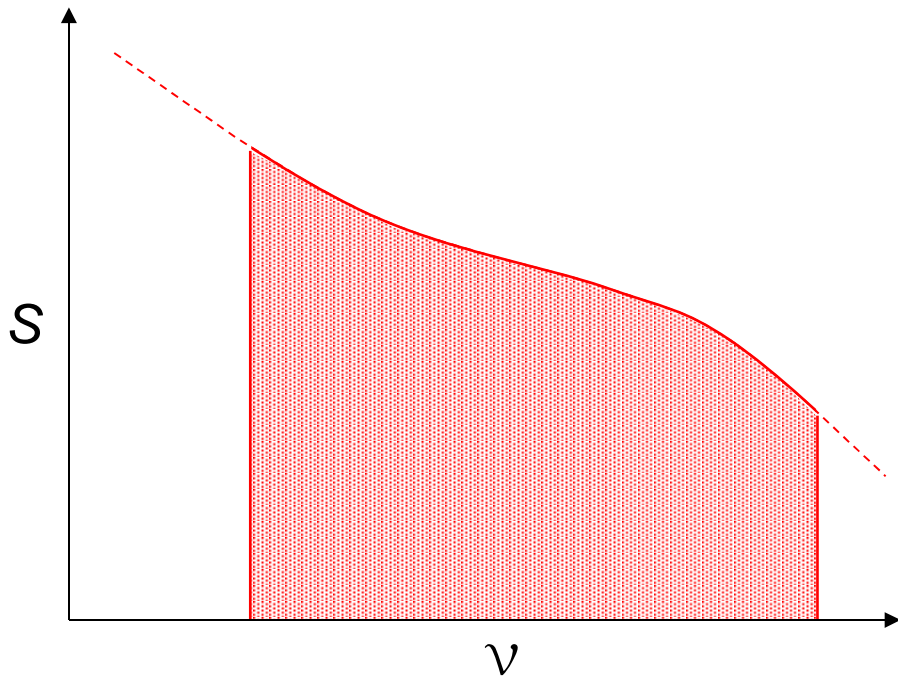
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To convince people that there is more to parallel CLEANing than Taylor expansion in frequency space.



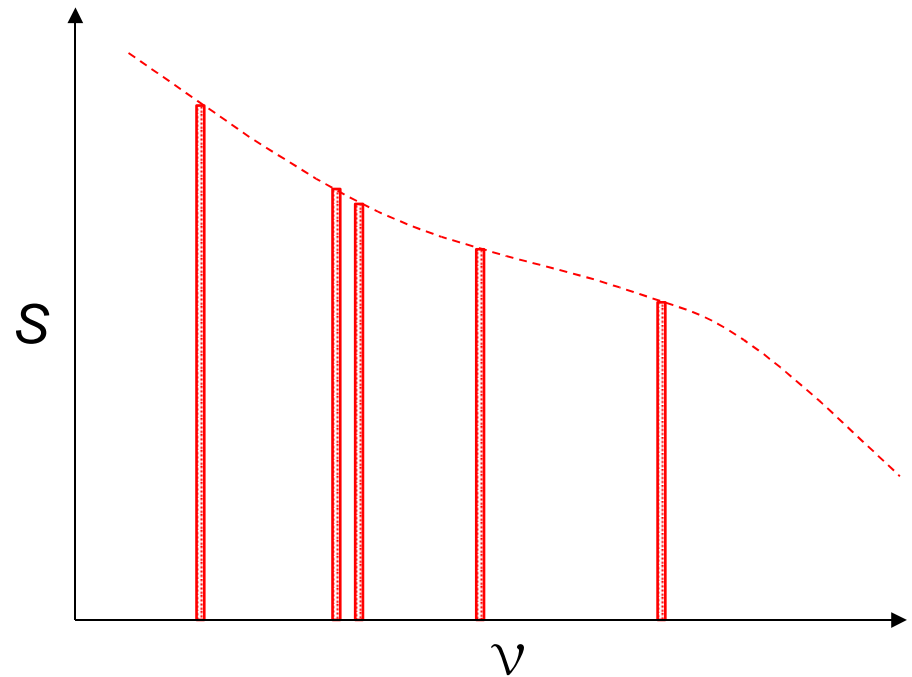
# MFS of several narrow-band epochs

Wide-band instruments:



Suitable for Sault-Wieringa  
(Taylor expansion)

Several narrow-band epochs:



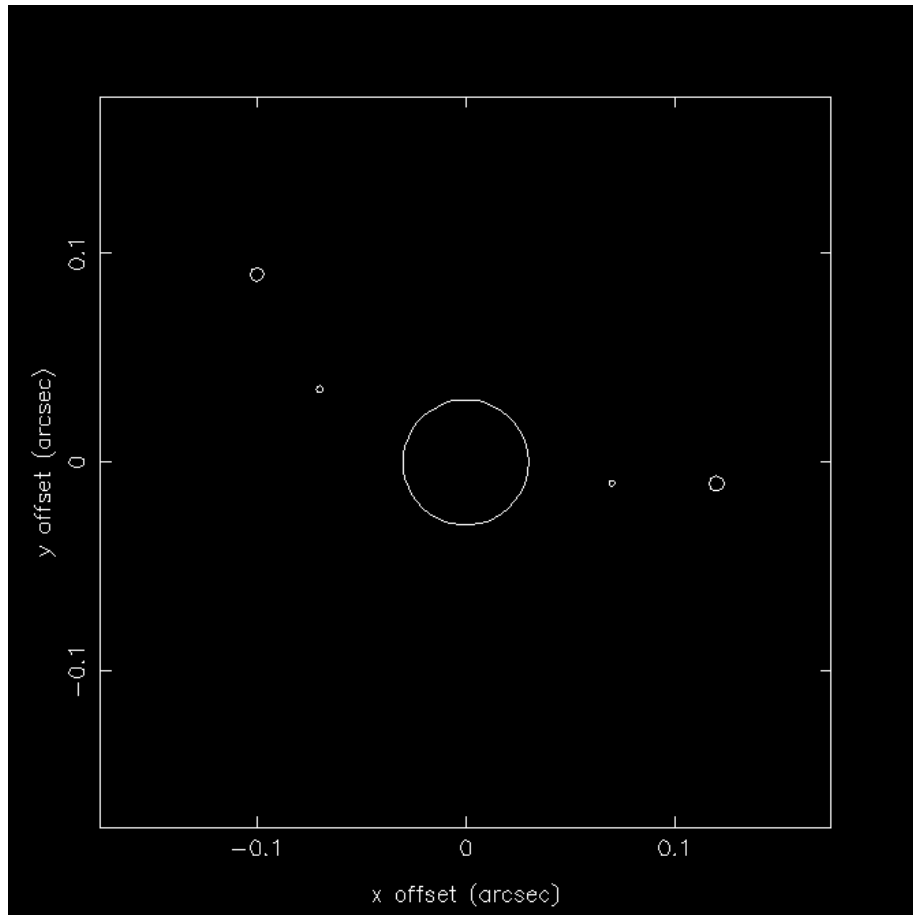
More natural:  
1 beam per epoch.

# Parallel CLEAN for time synthesis

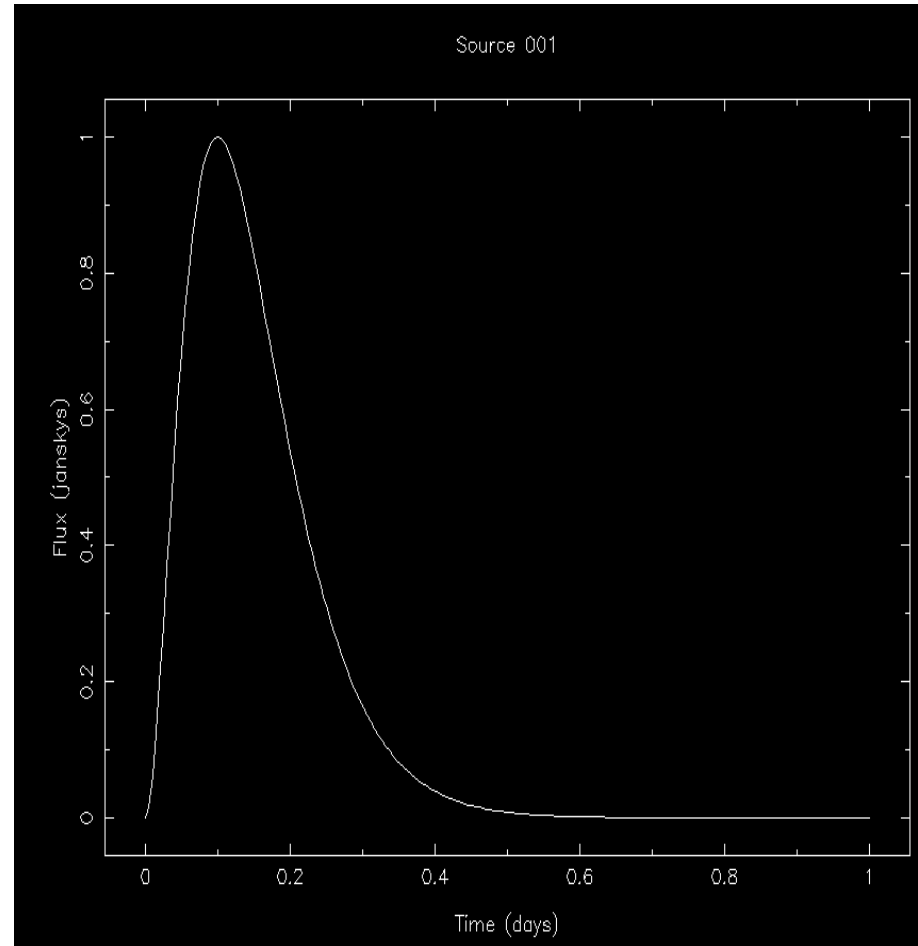
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# Parallel CLEAN for time synthesis

Model (time-averaged flux):



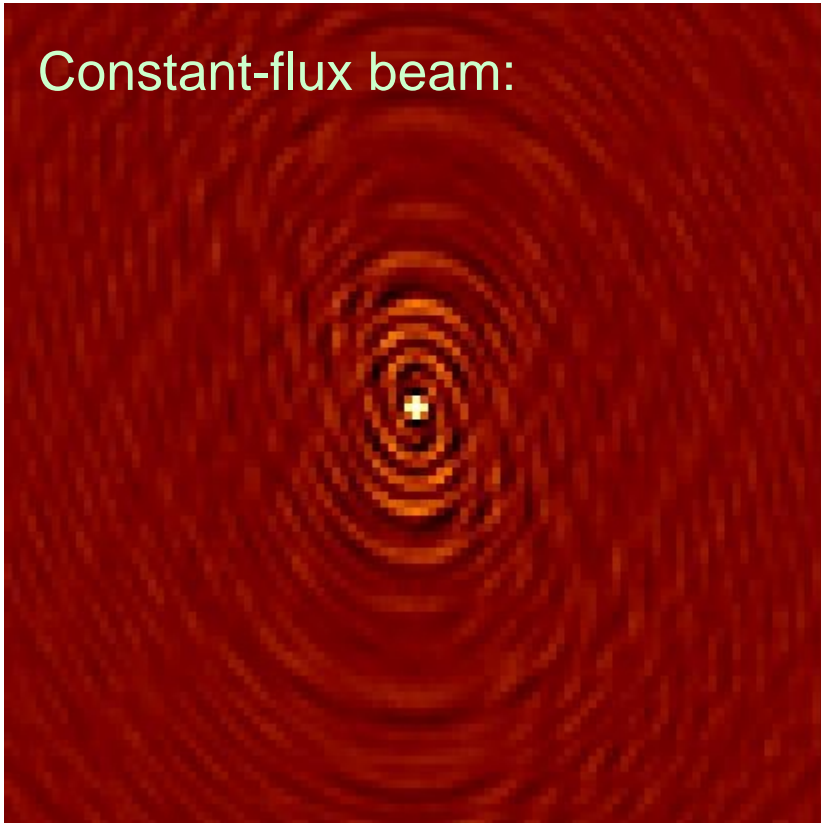
Light-curve of central source:



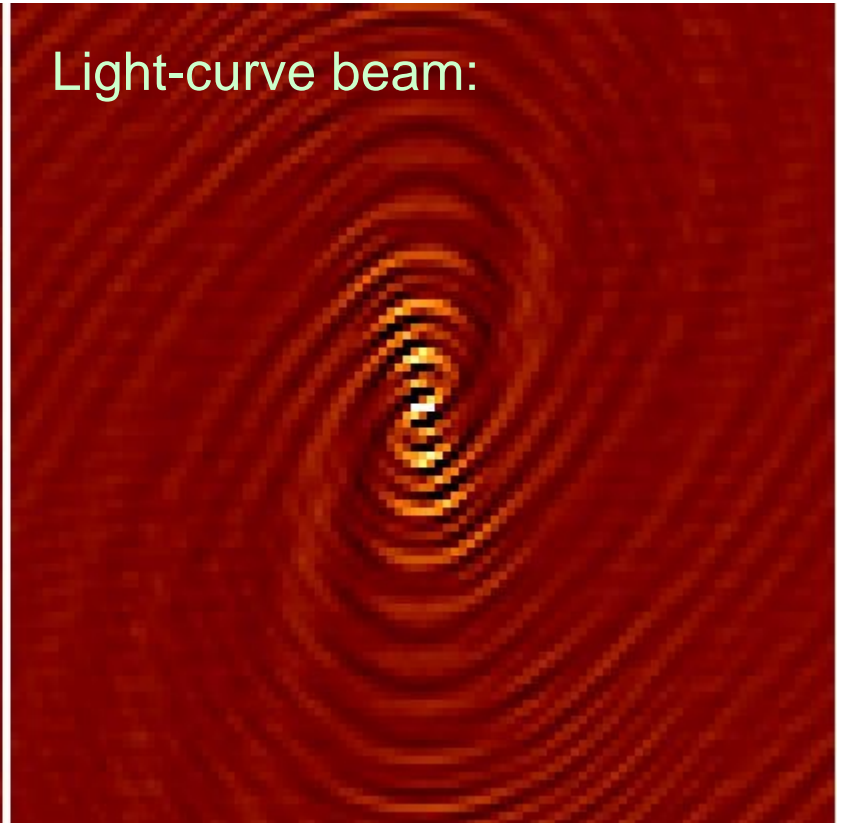
# Parallel CLEAN for time synthesis

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Constant-flux beam:



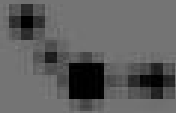
Light-curve beam:



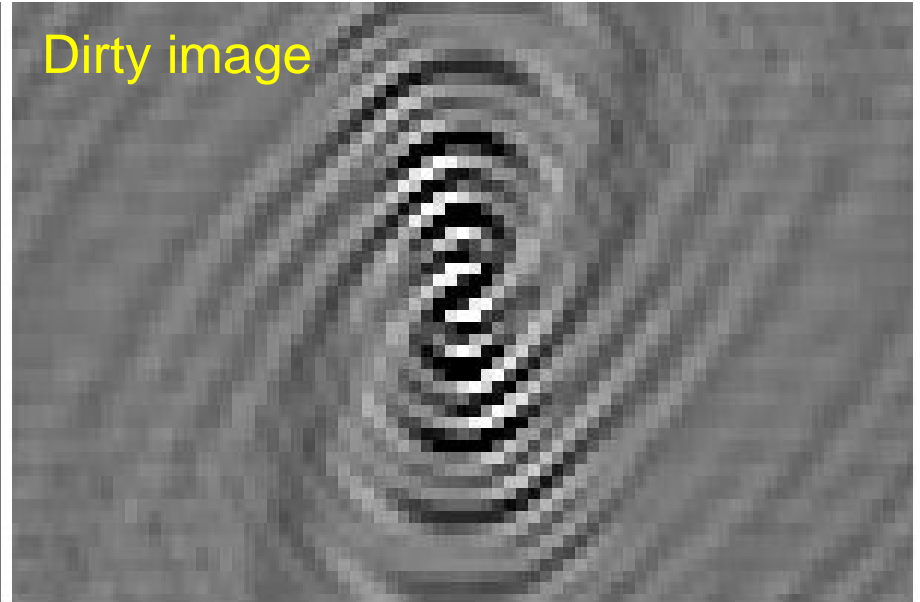
# Parallel CLEAN for time synthesis

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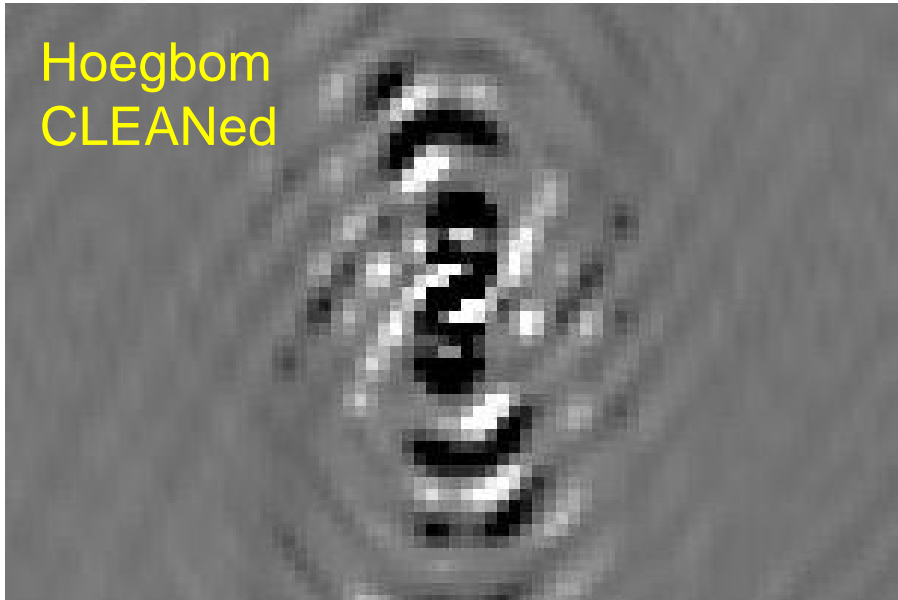
Model (conv. with RB)



Dirty image



Hoegbom  
CLEANed

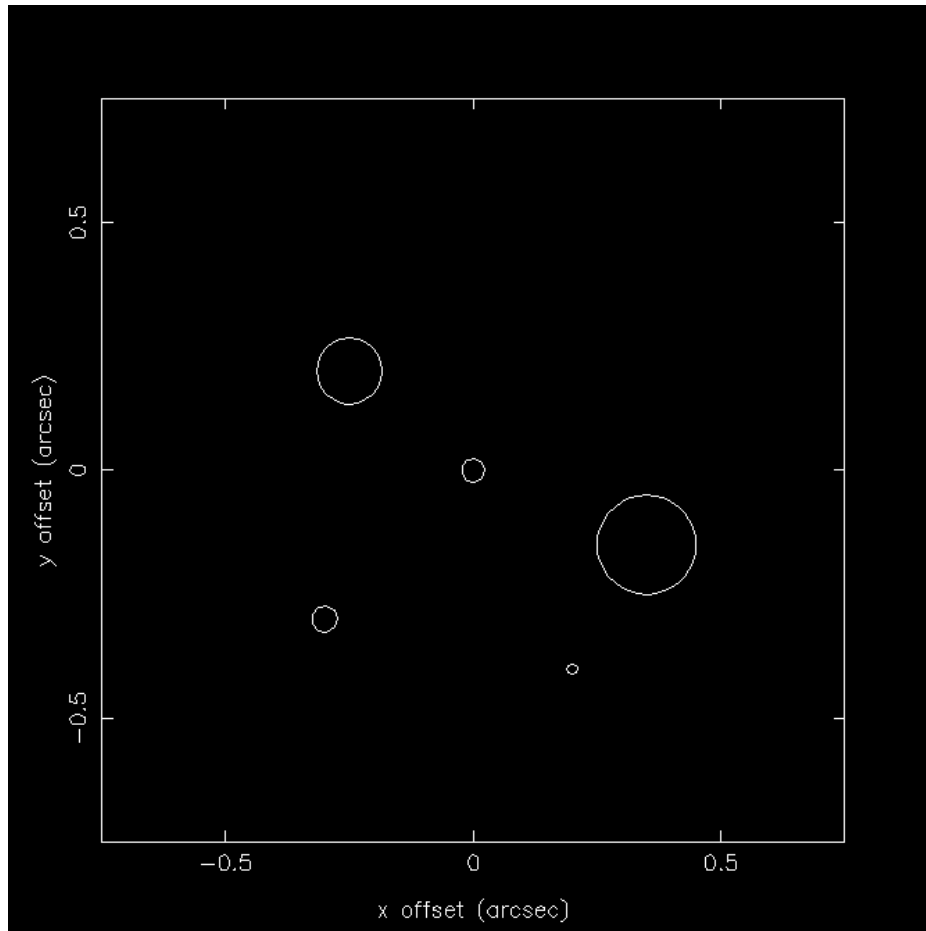


Parallel CLEANed

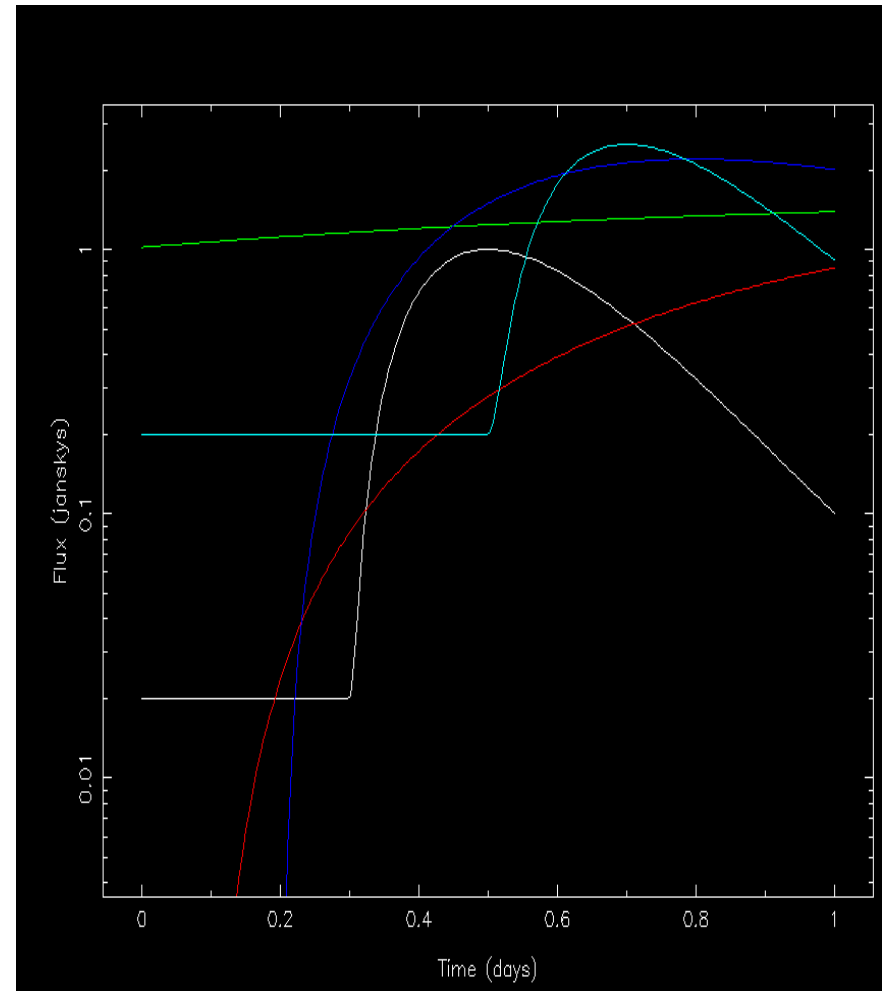


# Parallel CLEAN for time synthesis

Model (flux averaged over frequency and time):



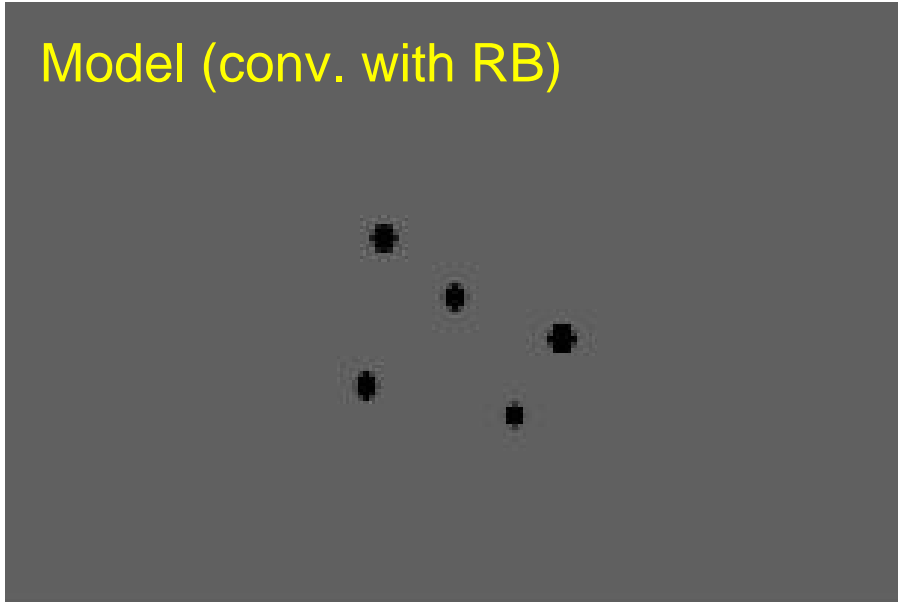
Light-curves:



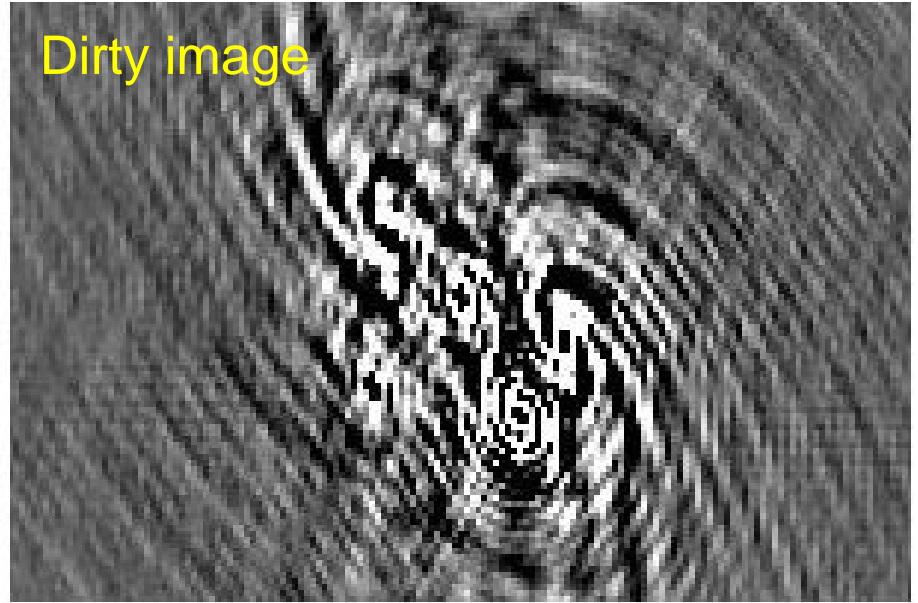


# Parallel CLEAN for time synthesis

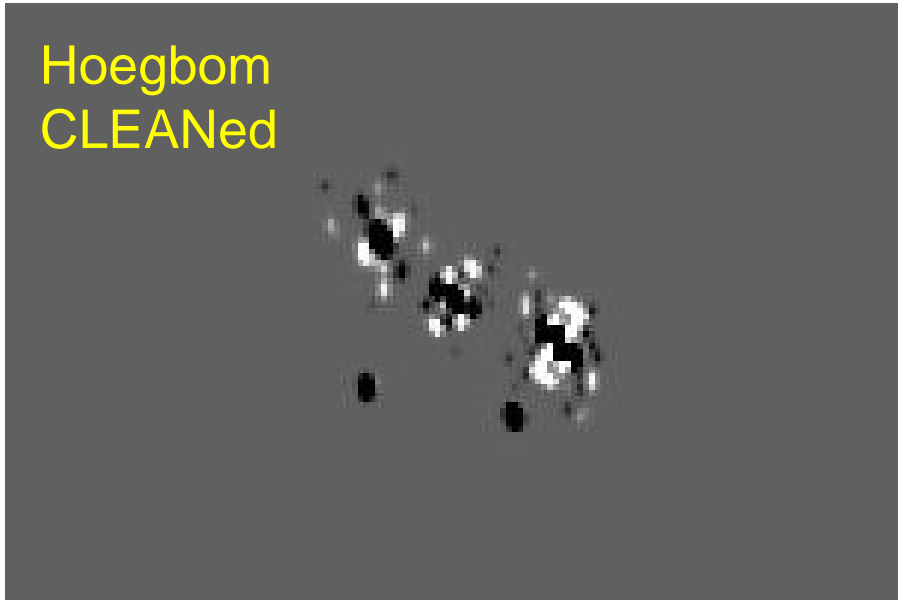
Model (conv. with RB)



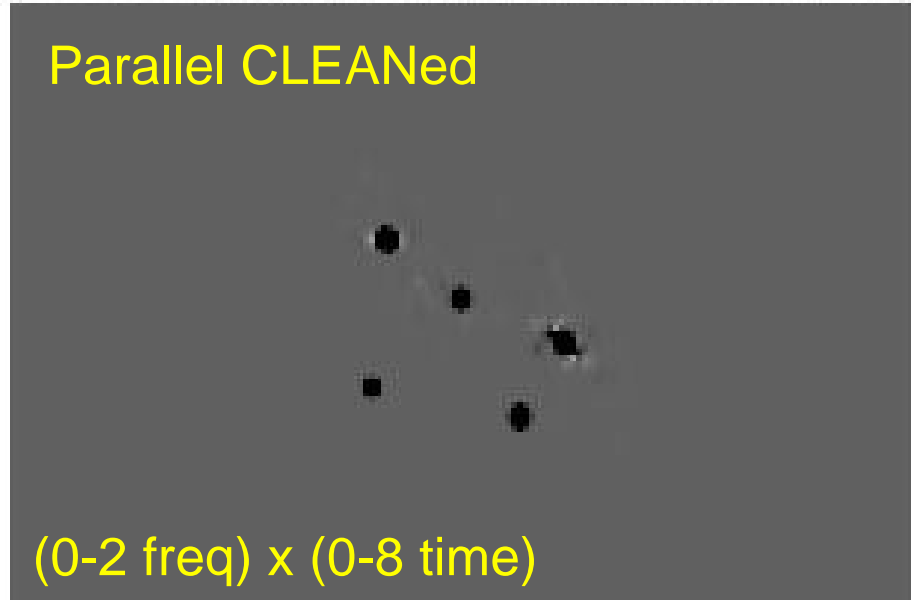
Dirty image



Hoegbom  
CLEANed



Parallel CLEANed



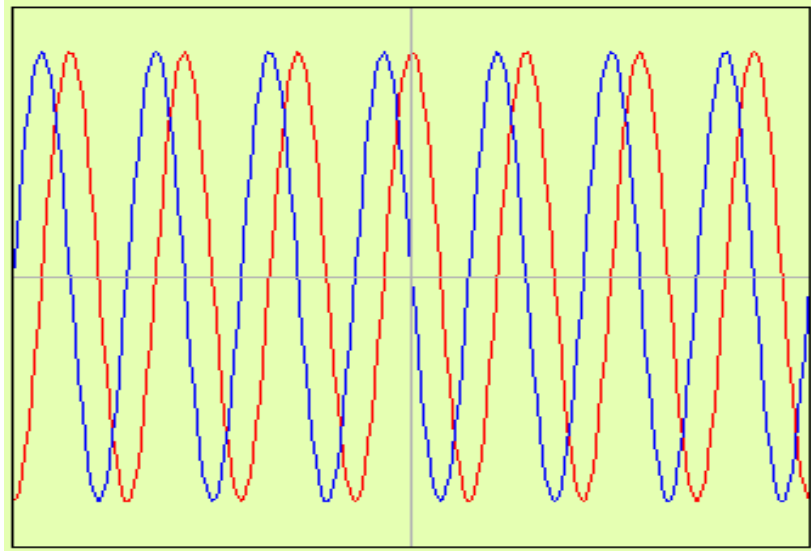
(0-2 freq) x (0-8 time)

# Off-pixel-centre sources

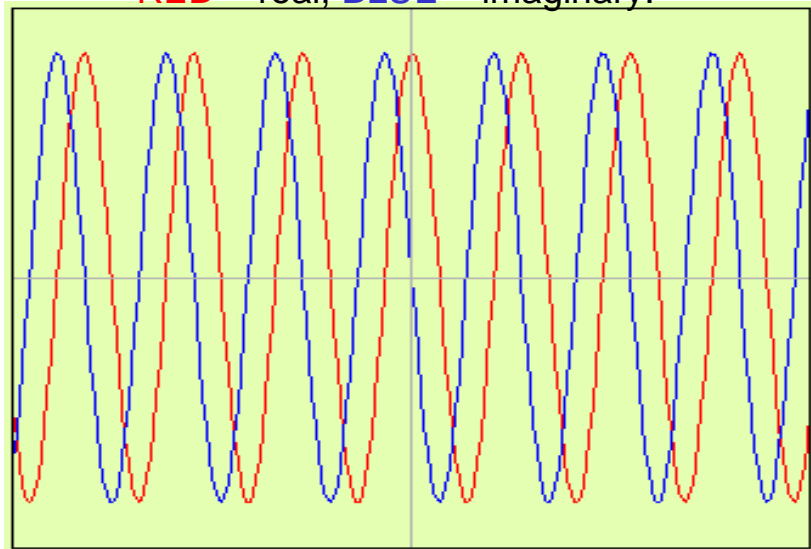
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# Off-pixel-centre sources

U-V 'plane'

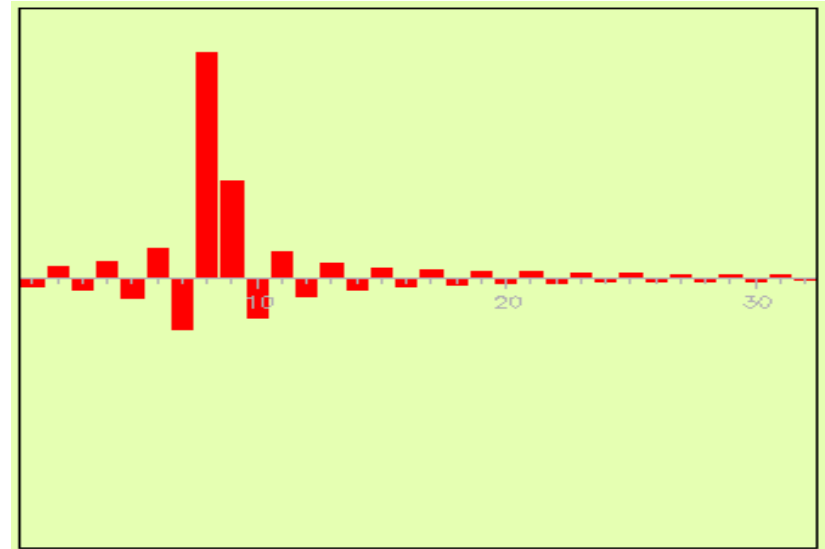
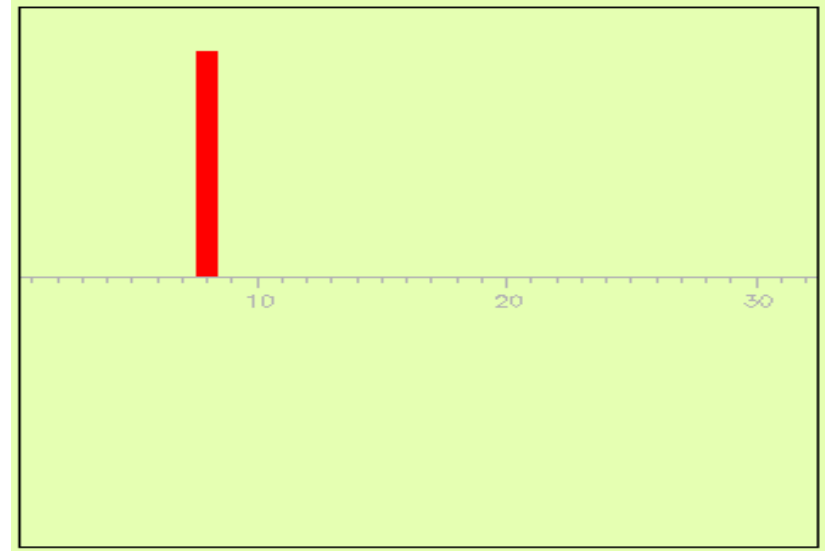


RED = real, BLUE = imaginary.



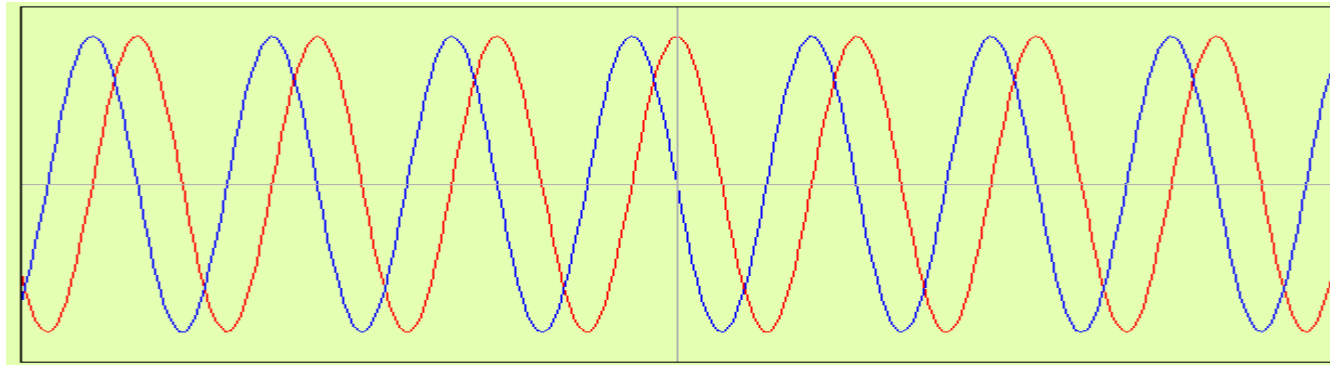
*Fourier*

'Sky'



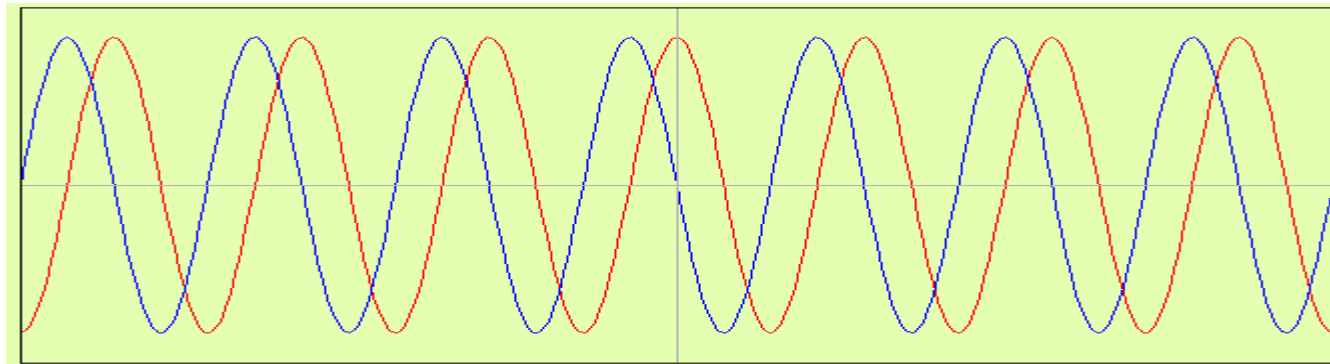
*Fourier*

# Off-pixel-centre sources



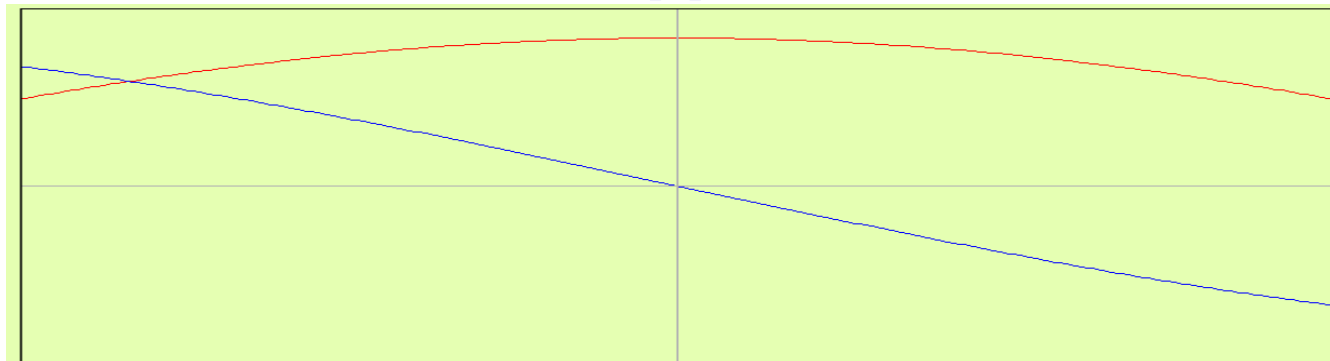
$$\exp\left(-2\pi i\left[n+f\right]\frac{t}{T}\right)$$

=



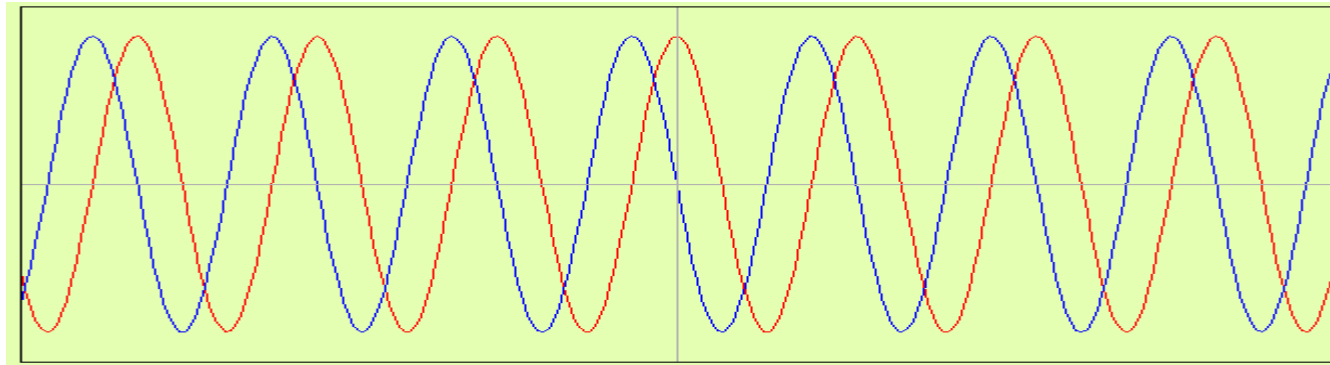
$$\exp\left(-2\pi i n \frac{t}{T}\right)$$

X



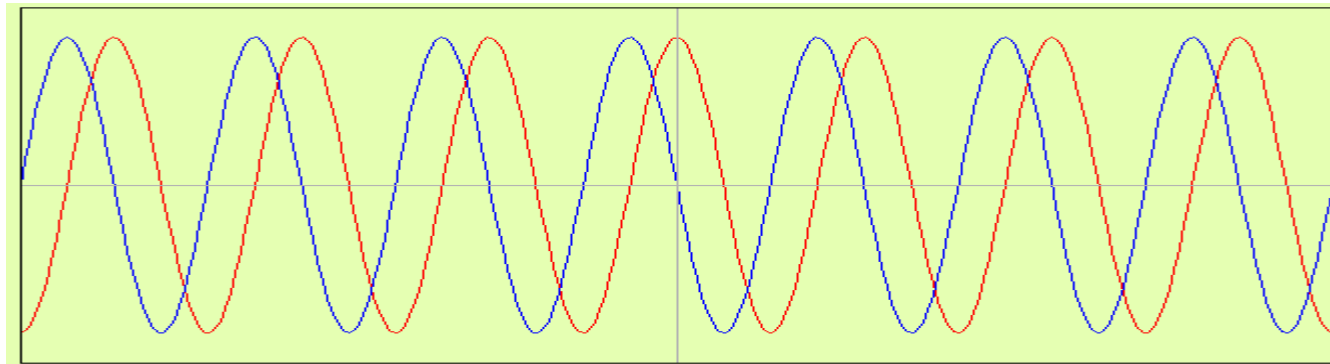
$$\exp\left(-2\pi i f \frac{t}{T}\right)$$

# Off-pixel-centre sources



$$\exp\left(-2\pi i\left[n+f\right]\frac{t}{T}\right)$$

~



$$\exp\left(-2\pi i n \frac{t}{T}\right)$$

X



$$1-2\pi i f \frac{t}{T}$$

# Off-pixel-centre sources

Lower left source at  $(m \Delta x - 0.4, n \Delta y + 0.2)$ :

