



Imaging and Calibration Algorithms for EVLA, e-MERLIN and ALMA

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- Issues are inextricably linked
- Primary beams (all Stokes)
 - Use a proper physical model
 - Need to do simulations (EM and FE) and measurements
 - Variations (El, T, ...) are not understood
 - Beam squint corrections understood
- Ionosphere
 - What is the most economical description?
 - MIM
 - Field-based calibration. Zernike basis or something else?
 - GPS data
 - Ionospheric physics





- Field-based calibration
- Peeling (various approaches)
- Global (expectation maximization)
- Empirical approaches (differential gain)
- Understanding the physics will lead to models with fewer parameters to be measured
- **Apply corruptions to the model and subtract**
- More use of smoothness in time or direction
- What are the errors?
- The "nicest" image is not necessarily the most faithful description of the sky.



Wide-field imaging approaches



Correcting for the w term

Facets (AIPS/Obit) with auto-boxing extensions
w-projection

Correction for direction-dependent effects
 Peeling + squint
 AW projection
 Pointing self-calibration



Deconvolution



Multi-scale clean approaches look promising:

- AIPS, CASA implementations
- ASP clean
- MEM

Discussion on convergence; different algorithms

Bayesian approach to point-source fitting, RM, etc.

Spectral-line issues

Calibration transfer between sub-bands
Very wide bandpasses
Multi-channel questions



Wide-band imaging



Sault-Wieringa to higher order. Can handle variable sources in similar ways

- MF MSC approach is a promising generalization
- Fitting cleaned narrow-band images is an alternative
- We do not have the equivalent for linear polarization

Need to integrate with wide-field imaging



What should we do next?



Document

- What problems do we need to solve? Comprehensive description using ME framework.
- What software is available?
- What is missing?
- SOC will write a framework and ask for input
- Wiki
 - Talks from this and other meetings
 - Papers, memos
 - Software documentation
 - Forum for interchange of ideas



Test problems



Standard test problems

- Simulated data without noise
- Simulated data with realistic noise
- Real data

Peeling and related algorithms

- Flavours of peel
- Expectation maximization

Wide-band imaging

- Sault-Wieringa to various orders
- Hybrid
- MS-MSC

Fitting deconvolved narrow-band images



Test problems (2)



Wide-field imaging

- w-projection
- facets
- Deconvolution
 - standard clean
 - multiscale clean variants
 - MEM variants



Promoting collaboration



Essential to document algorithms in detail

- Publish, but also
- provide code/detailed description
- Funding for visits
- More meetings
 - how often?
 - who?

Proposals for additional funding?

- Observatories how to approach?
- PhD projects
- National/International funding agencies







- We have many of the algorithms we need, but not all
- Terabyte data problem
- We have made progress on interoperability at the scripting level (Parseltongue, python interfaces to Obit and CASA)
- It is very inefficient to exchange data between CASA and AIPS/Obit data: we cannot afford to do this more than once or twice in the data-reduction process.
- AIPS/Obit and CASA calibration models are very different, so need to apply calibrations to data before interchanging.
- How do we proceed?