

# Expectation Maximization Type Algorithms for Calibration

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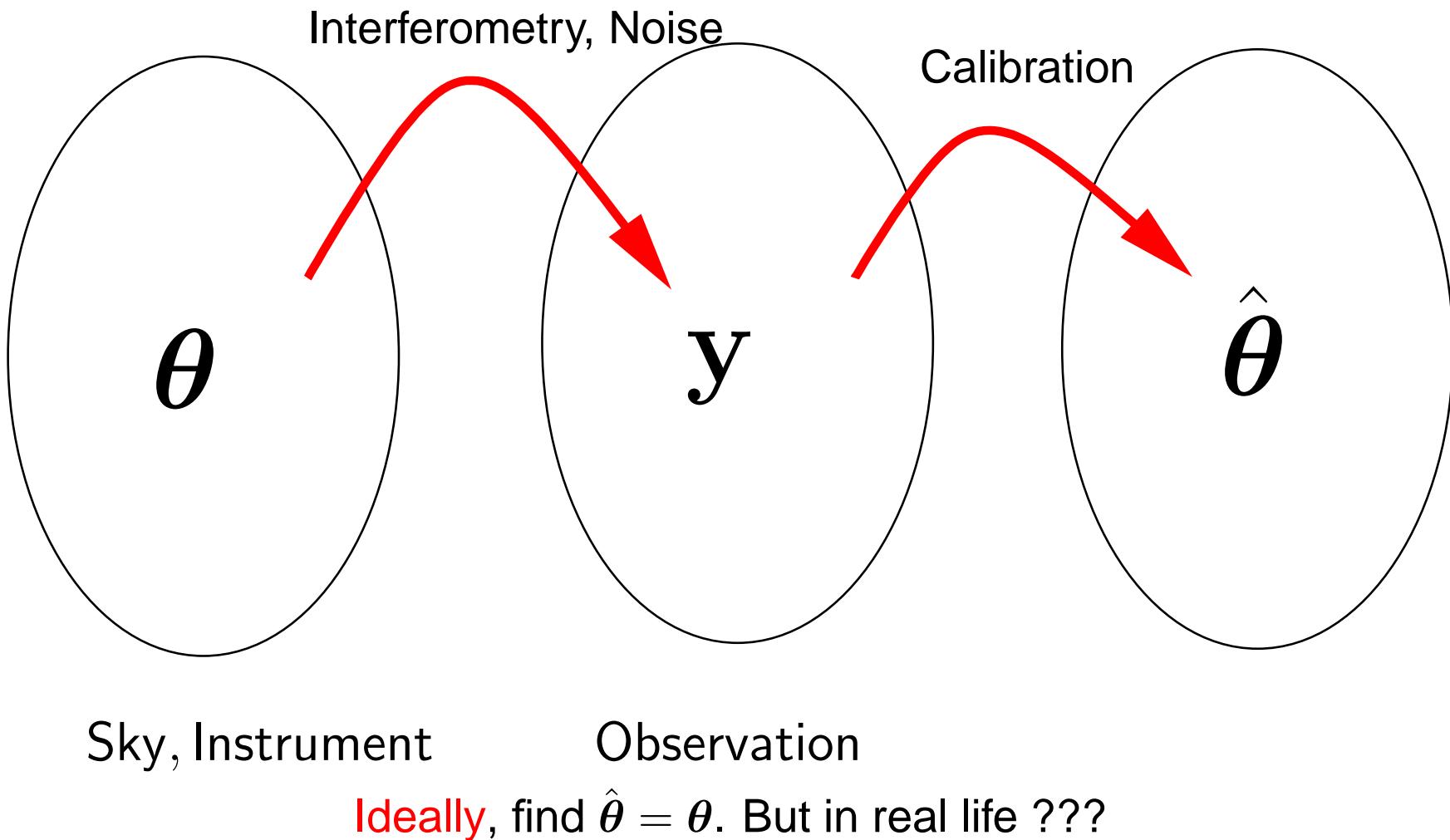
and

ASTRON

collaborators:

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# Calibration



# Calibration

For  $K$  discrete sources, we observe

$$\mathbf{y} = \sum_{i=1}^K \mathbf{s}_i(\boldsymbol{\theta}) + \mathbf{n}$$

Maximum Likelihood (ML) estimate, under White Gaussian Noise

$$\hat{\boldsymbol{\theta}} = \arg \min_{\boldsymbol{\theta}} \phi(\boldsymbol{\theta}) = \arg \min_{\boldsymbol{\theta}} \|\mathbf{y} - \sum_{i=1}^K \mathbf{s}_i(\boldsymbol{\theta})\|^2$$

Traditional calibration: using Levenberg-Marquardt (LM) algorithm

$$\boldsymbol{\theta}^{k+1} = \boldsymbol{\theta}^k - (\nabla_{\boldsymbol{\theta}} \nabla_{\boldsymbol{\theta}}^T \phi(\boldsymbol{\theta}) + \lambda \mathbf{H})^{-1} \nabla_{\boldsymbol{\theta}} \phi(\boldsymbol{\theta})|_{\boldsymbol{\theta}^k}$$

where  $\mathbf{H} \triangleq \text{diag}(\nabla_{\boldsymbol{\theta}} \nabla_{\boldsymbol{\theta}}^T \phi(\boldsymbol{\theta}))$ .

# LM in Action



The long and winding road

# Expectation Maximization

- Formally introduced by [Dempster, Laird, Rubin, 77]
- ML estimation with incomplete data
- Iterative maximization of the likelihood
- Frequently used but most misunderstood algorithm

E.g. Gain/Leakage calibration with an unpolarized calibrator

$$\mathbf{V}_{pq} = I\mathbf{G}_p\mathbf{D}_p\mathbf{D}_q^H\mathbf{G}_q^H + \mathbf{N}$$

$$\mathbf{G}_p = \begin{bmatrix} g_{x,p} & 0 \\ 0 & g_{y,p} \end{bmatrix}, \quad \mathbf{D}_p = \begin{bmatrix} 1 & \delta_{x,p} \\ \delta_{y,p} & 1 \end{bmatrix}, \quad \delta_{x,p}, \delta_{y,p} \ll 1$$

First Estimate  $\mathbf{G}$  and then  $\mathbf{D}$ , and iterate.  
But, does the other way work?

# EM: Formal Description

$$\mathbf{y} = \sum_{i=1}^K \mathbf{s}_i(\boldsymbol{\theta}) + \mathbf{n}$$

- ML estimate:  $\hat{\boldsymbol{\theta}}_{ML} = \arg \max_{\boldsymbol{\theta}} \log f(\mathbf{y}|\boldsymbol{\theta})$
- Auxiliary random variable  $\mathbf{x}$ : hidden data,  $\mathbf{y} = \mathbf{F}(\mathbf{x})$
- The *E Step*: compute conditional expectation  
$$Q(\boldsymbol{\theta}|\boldsymbol{\theta}^k) = E\{\log f(\mathbf{x}|\boldsymbol{\theta})|\mathbf{y}, \boldsymbol{\theta}^k\}$$
- The *M Step*: Maximize  $\boldsymbol{\theta}^{k+1} = \arg \max_{\boldsymbol{\theta}} Q(\boldsymbol{\theta}|\boldsymbol{\theta}^k)$
- Can be simplified for exponential family distributions.
- Can be even more simplified for Gaussian distributions.

# Classic EM

- Auxiliary random variables

$$\tilde{\mathbf{x}}_i = \mathbf{s}_i(\boldsymbol{\theta}_i) + \tilde{\mathbf{n}}_i$$

- Noise

$$\mathbf{n} = \sum_{i=1}^K \tilde{\mathbf{n}}_i, \quad E\{\tilde{\mathbf{n}}_i \tilde{\mathbf{n}}_j^H\} = \beta_i \delta_{ij} \boldsymbol{\Pi}, \quad \sum_{i=1}^K \beta_i = 1$$

- *E Step:*

$$\hat{\tilde{\mathbf{x}}}_i = \mathbf{s}_i(\boldsymbol{\theta}_i^k) + \beta_i (\mathbf{y} - \sum_{l=1}^K \mathbf{s}_l(\boldsymbol{\theta}_l^k))$$

- *M Step:*

$$\boldsymbol{\theta}_i^{k+1} = \boldsymbol{\theta}_i^k - (\nabla_{\boldsymbol{\theta}_i} \nabla_{\boldsymbol{\theta}_i}^T \phi_i(\boldsymbol{\theta}_i) + \lambda \mathbf{H}_i)^{-1} \nabla_{\boldsymbol{\theta}_i} \phi_i(\boldsymbol{\theta}_i)|_{\boldsymbol{\theta}_i^k}$$

# SAGE

SAGE: Space Alternating Generalized Expectation Maximization [Fessler and Hero, 94]

- Auxiliary random variable

$$\mathbf{x}^S = \mathbf{s}_i(\boldsymbol{\theta}_i) + \mathbf{n}$$

- *E Step:*

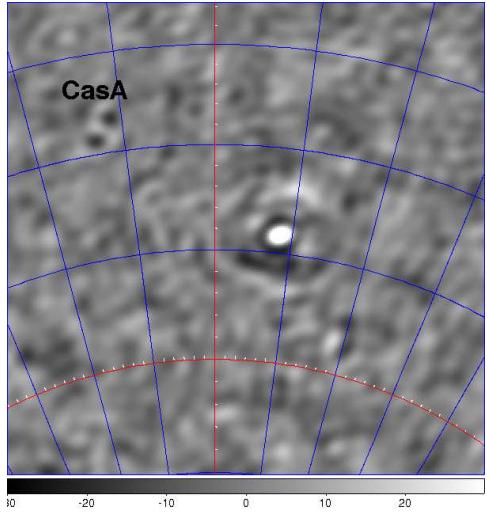
$$\widehat{\mathbf{x}^S} = \mathbf{s}_i(\boldsymbol{\theta}_i^k) + (\mathbf{y} - \sum_{l=1}^K \mathbf{s}_l(\boldsymbol{\theta}_l^k)) = \mathbf{y} - \sum_{l=1, l \neq i}^K \mathbf{s}_l(\boldsymbol{\theta}_l^k)$$

- *M Step:*

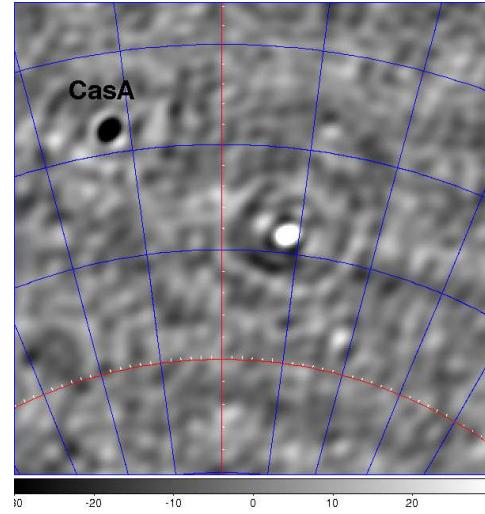
$$\boldsymbol{\theta}_i^{k+1} = \boldsymbol{\theta}_i^k - (\nabla_{\boldsymbol{\theta}_i} \nabla_{\boldsymbol{\theta}_i}^T \phi_i(\boldsymbol{\theta}_i) + \lambda \mathbf{H}_i)^{-1} \nabla_{\boldsymbol{\theta}_i} \phi_i(\boldsymbol{\theta}_i) |_{\boldsymbol{\theta}_i^k}$$

- Caveat  $f(\mathbf{y}, \mathbf{x}^S | \boldsymbol{\theta}) = f(\mathbf{y} | \mathbf{x}^S, \boldsymbol{\theta}_{\tilde{S}}) f(\mathbf{x}^S | \boldsymbol{\theta})$
- Faster convergence than the classic EM

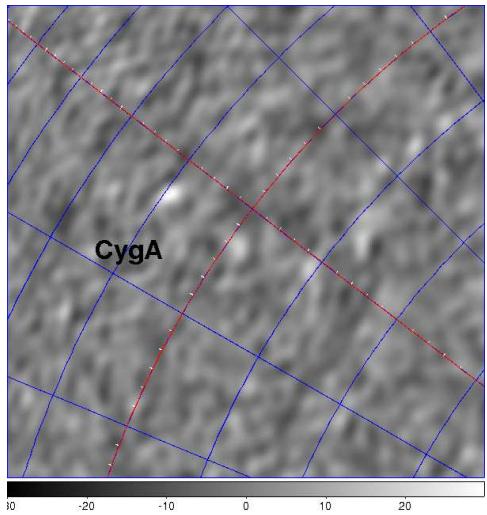
# Example-1 (LOFAR)



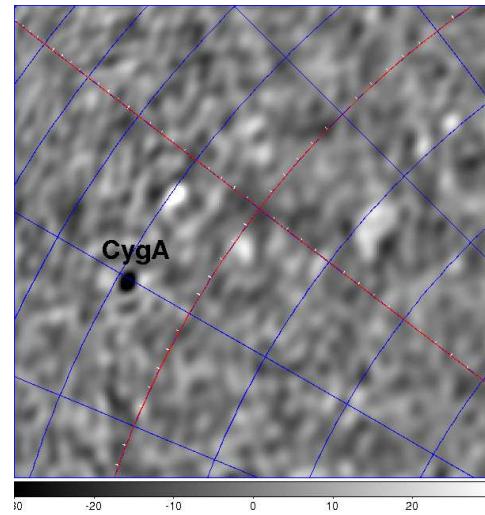
CasA, SAGE Algorithm



CasA, Normal Algorithm

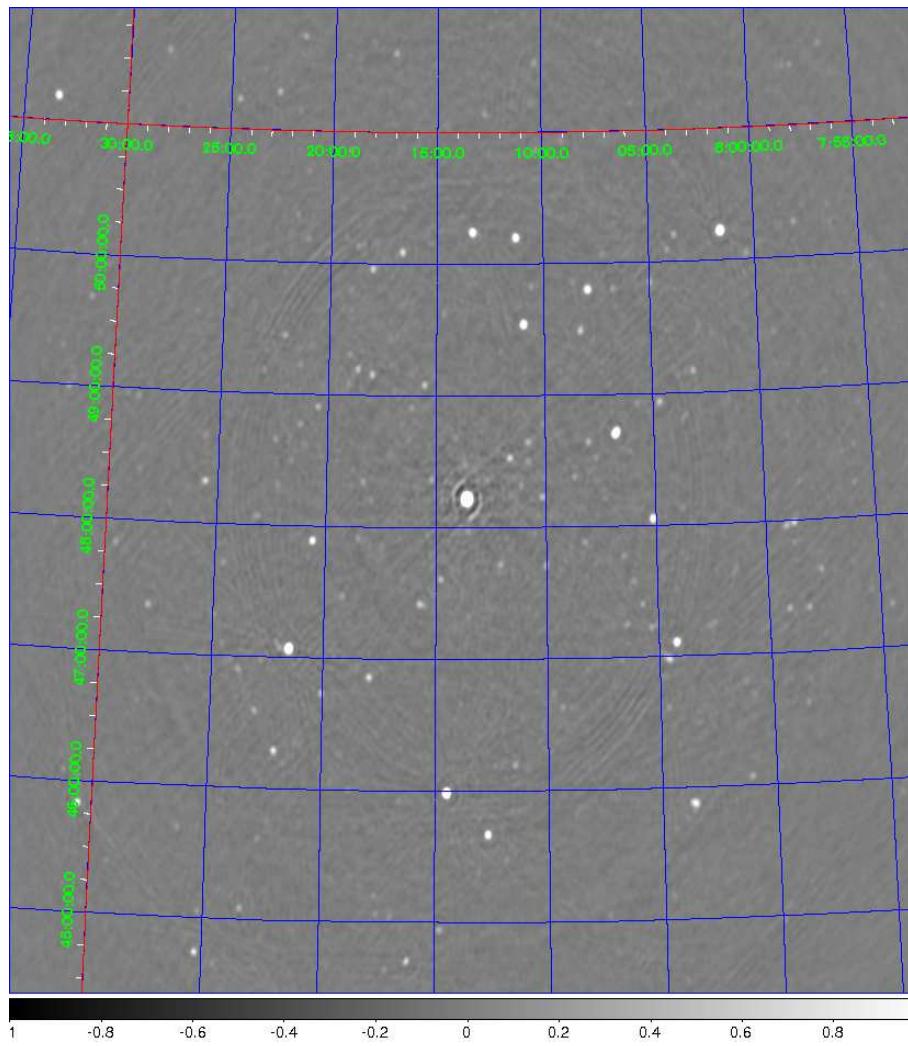


CygA, SAGE Algorithm  
ASTRON

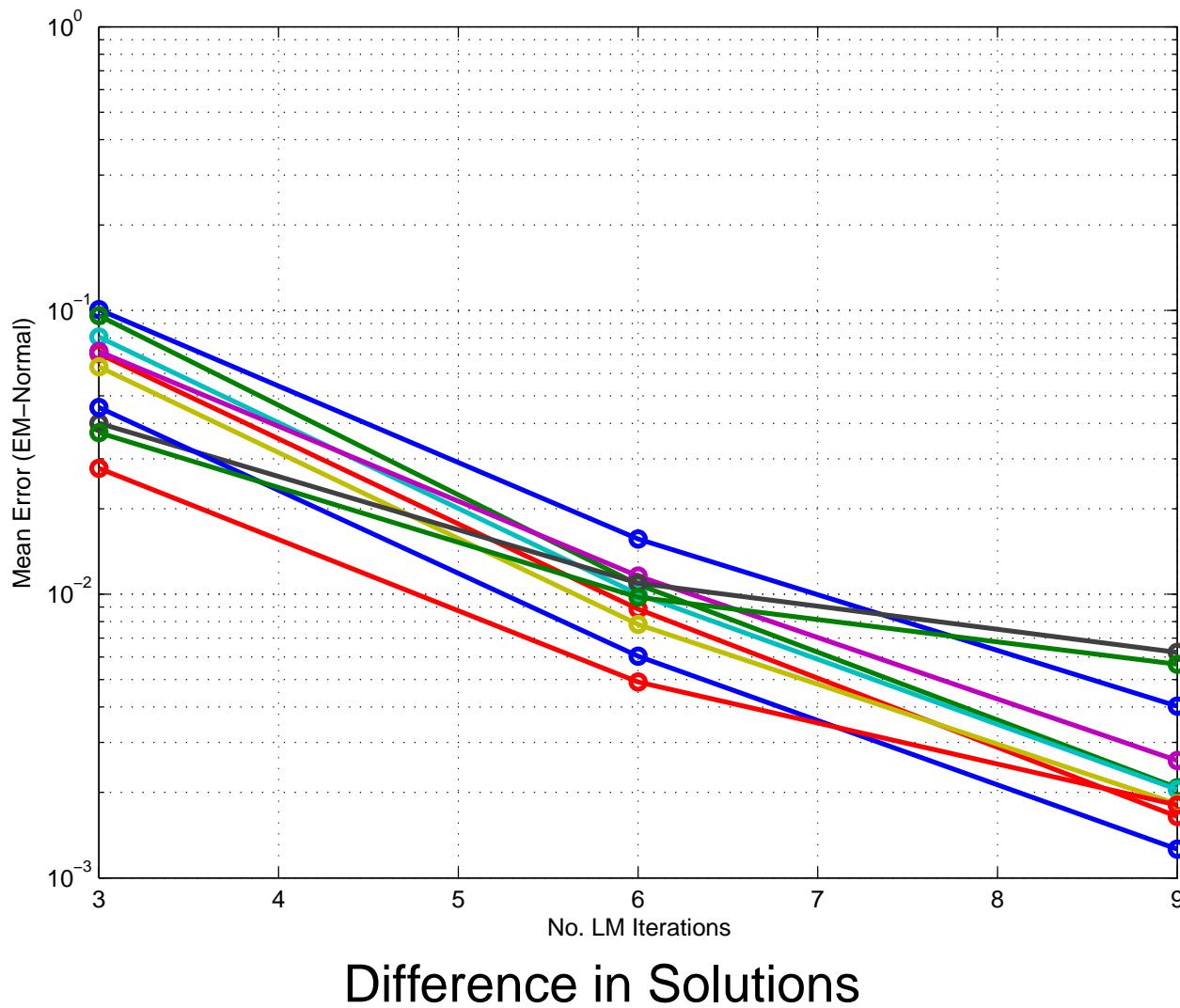


CygA, Normal Algorithm

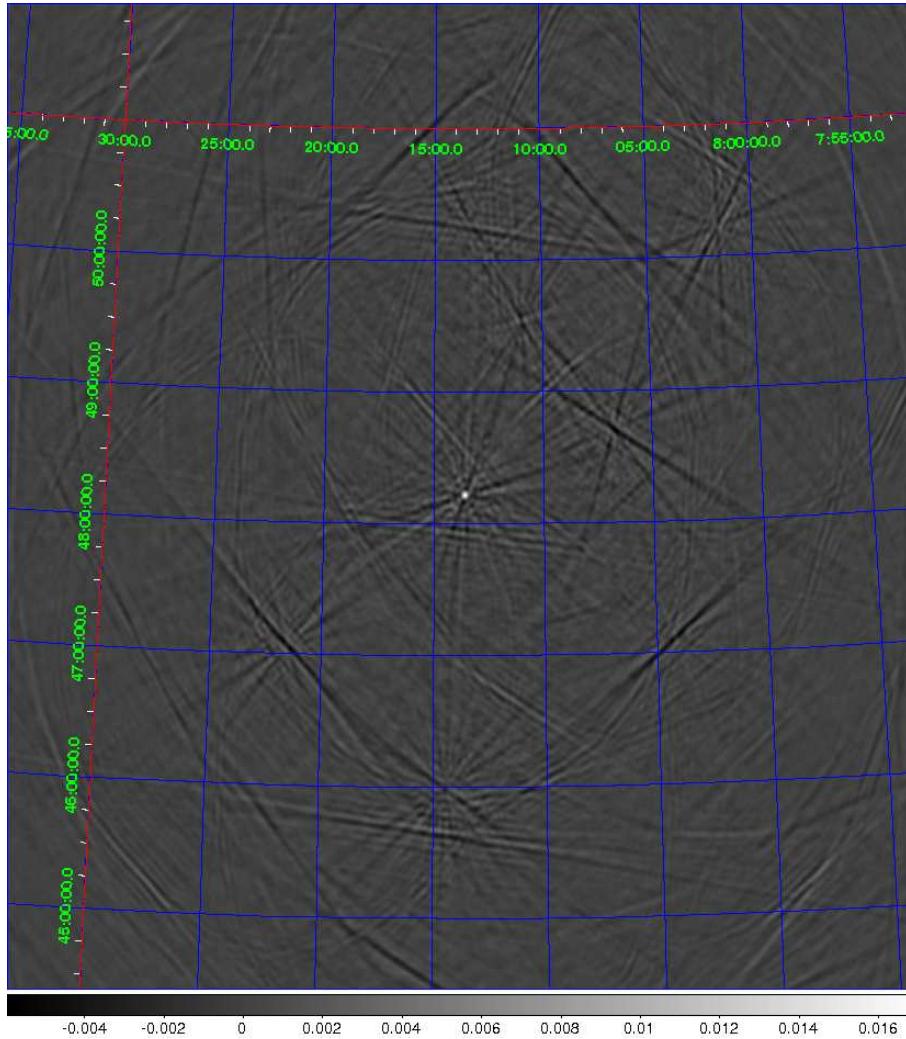
# Example-2



# Example-2



# Example-2



Difference in Normal and SAGE algorithms

# Conclusions

- EM/SAGE can reduce computational cost from  $\mathbb{O}(K^2N^2)$  to  $K\mathbb{O}(N^2)$  for  $KN$  parameters
- Can also improve quality of solutions when not converging
- See [Yatawatta,Zaroubi et al. 2008] for more information
- There is plenty more work to be done
- Advertisement: We are looking for a PhD student